### ARTE

**Astronomical Radio Transients Experiment** 

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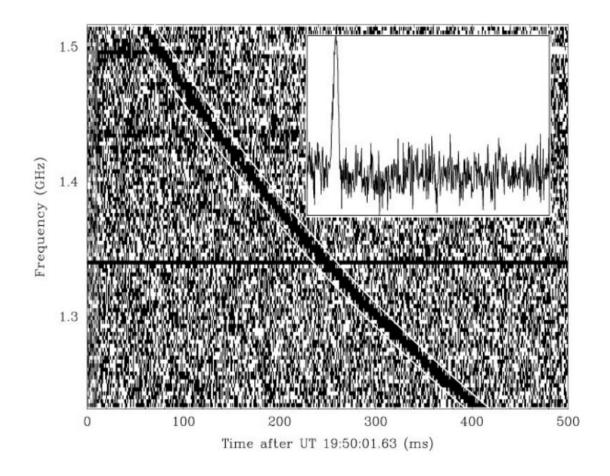


### Context

#### **FRBs**

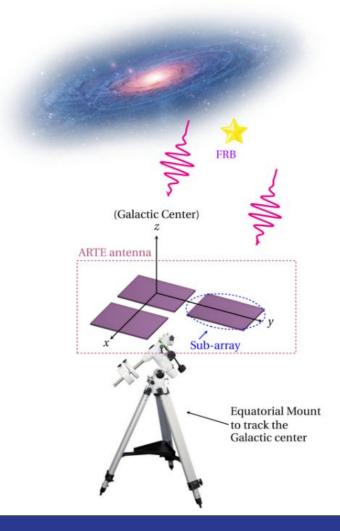
#### **Fast Radio Bursts**

- Short-duration, very high-energy radio wave pulses.
- Origin largely unknown.
- They are dispersed by propagation through ionized material in the interstellar and intergalactic medium.



#### **ARTE**

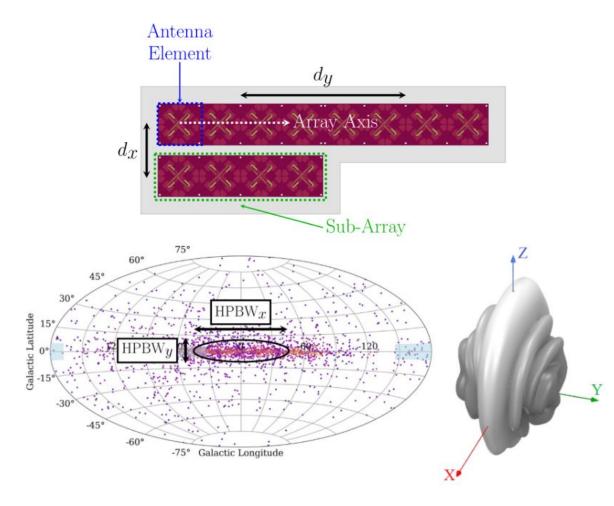
- Radio telescope located on Cerro Calán at the National Astronomical Observatory.
- It seeks to detect FRBs from the Galactic Center.
- Designed, built, and operated entirely by the team.
- Operates between 1.2 and 1.8 GHz.



## Antenna

#### **Radiation Pattern**

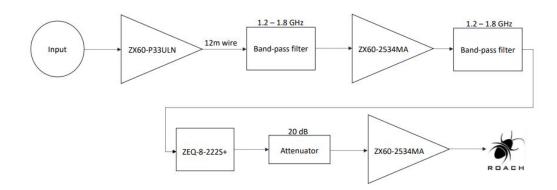
- There are 3
   arrangements
   distributed in an L
   shape.
- The antenna beam is designed to match the angular distribution of the center of the Milky Way.
- Maximizes gain in the center of the Milky Way.



## Microwave Receiver

#### Microwave Receiver

- A receiver is required for each antenna.
- It is mainly composed of amplifiers and filters.

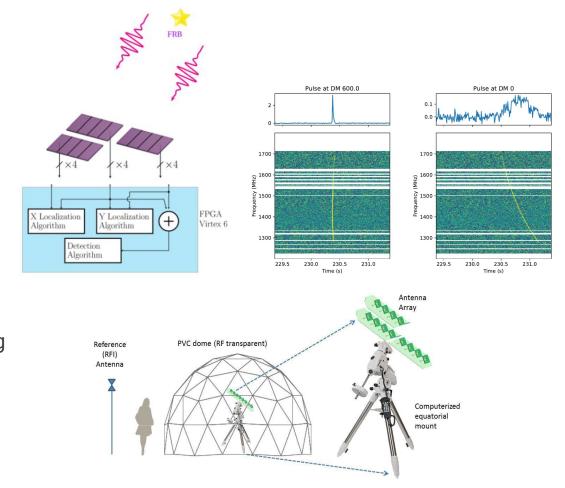




## Digital System

#### **Principal Features**

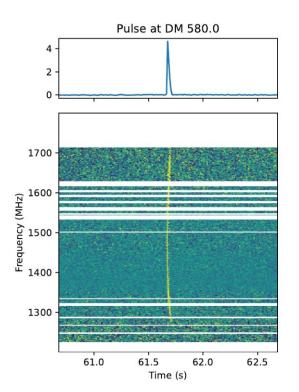
- DoA (Direction of Arrival algorithm): L-shaped configuration of the antennas allows us to determine the direction of arrival of the FRB.
- We installed a fourth omnidirectional antenna that detects ambient signals allowing us to discard radio noise.
- Real time detection using dedispersion.

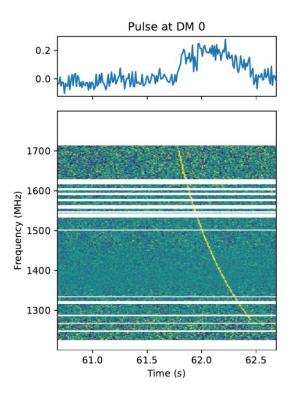


# Data Processing and Detections

#### Post-processing

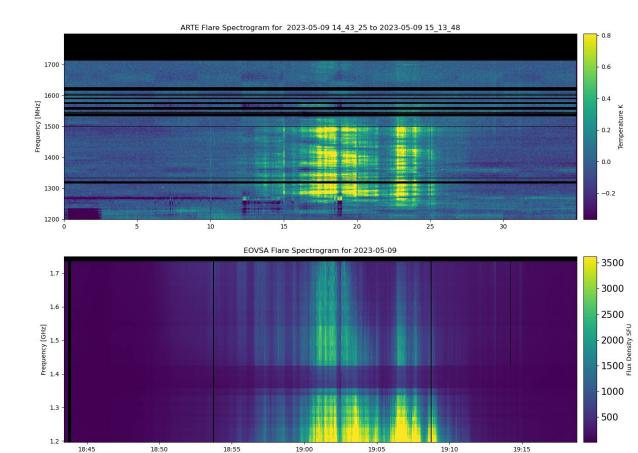
 Post-processing of the data is performed with a custom-developed pipeline that utilizes PRESTO, a software specifically designed for radio pulse detection.





#### Solar Flare

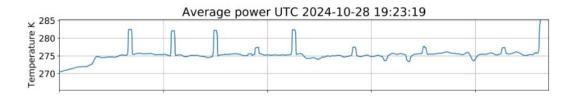
- Example of a solar flare detected by ARTE in May 2023.
- Comparison with data from EOVSA, a radio telescope dedicated to studying the Sun between frequencies of 1 and 18 GHz.

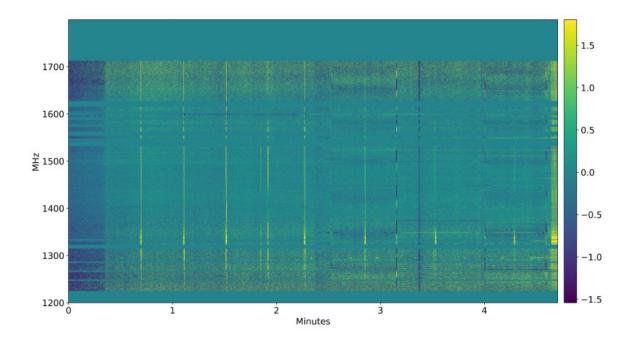


Time [UTC]

#### Synthetic FRBs

Using an RFSoC (new generation of FPGAs), we generated pulses simulating FRBs. These are injected directly into the ARTE receiver or triggered from a nearby antenna.





## **Future Work**

#### **Future Work**

- Achieve 100% automatic operation.
- Implement a raw data dump when a pulse is detected in real-time.
- Implement the second polarization.



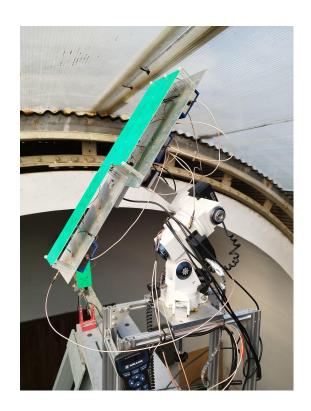
## Thank you

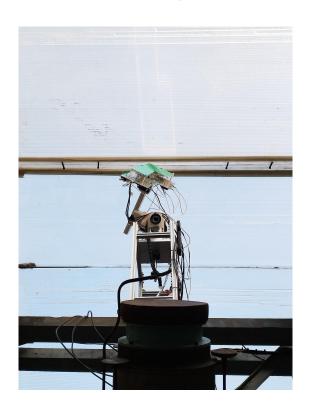
clsanmar@das.uchile.cl





### Antenna installed in the Meridiano building





## Receiver and Digital System installed in the Meridiano building



