

Parte I

Procesos físicos

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1. Consideraciones generales

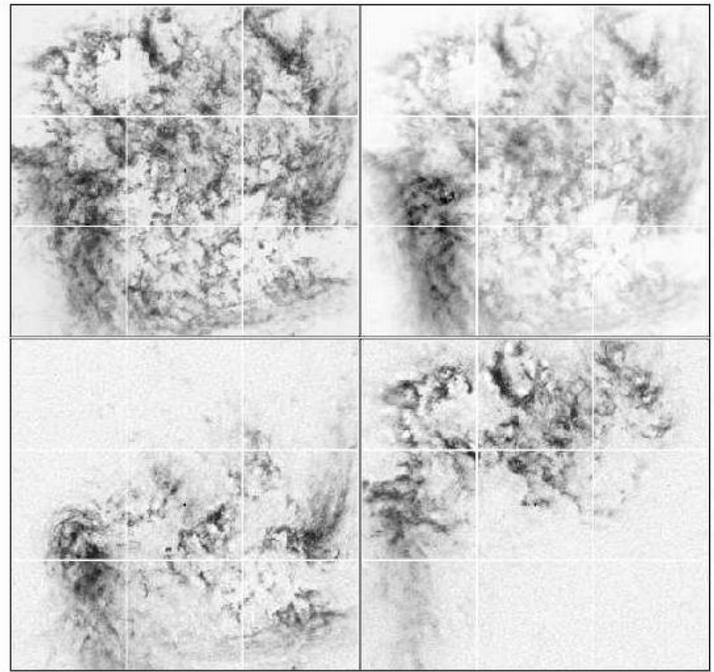
1.1. Equilibrio hidroestático

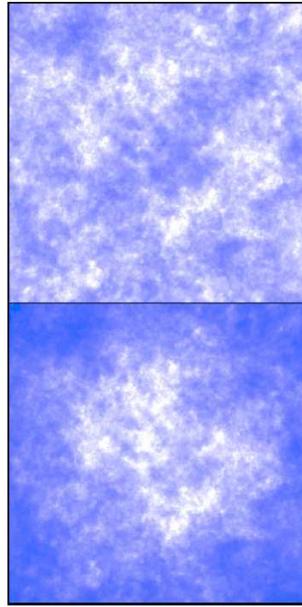
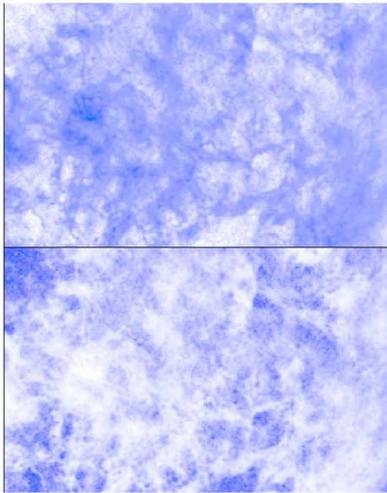
Teorema del virial

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1.2. Formación estelar

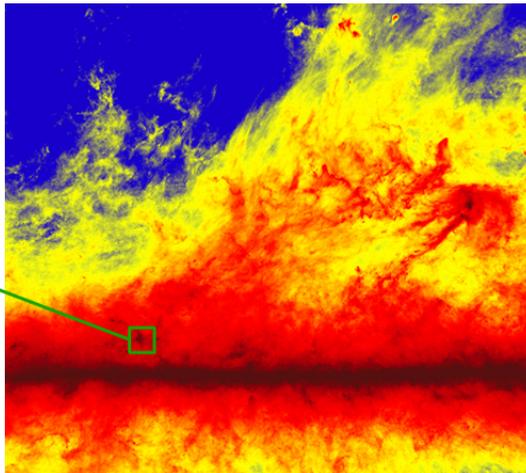
Análisis fractal de HI 21 cm en LMC (Elmegreen, Kim, Staveley-Smith ,2001,ApJ,548,749).





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Example star forming region: Aquila Rift



Left: Herschel, Right: Planck

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Example star forming region: Coal sack

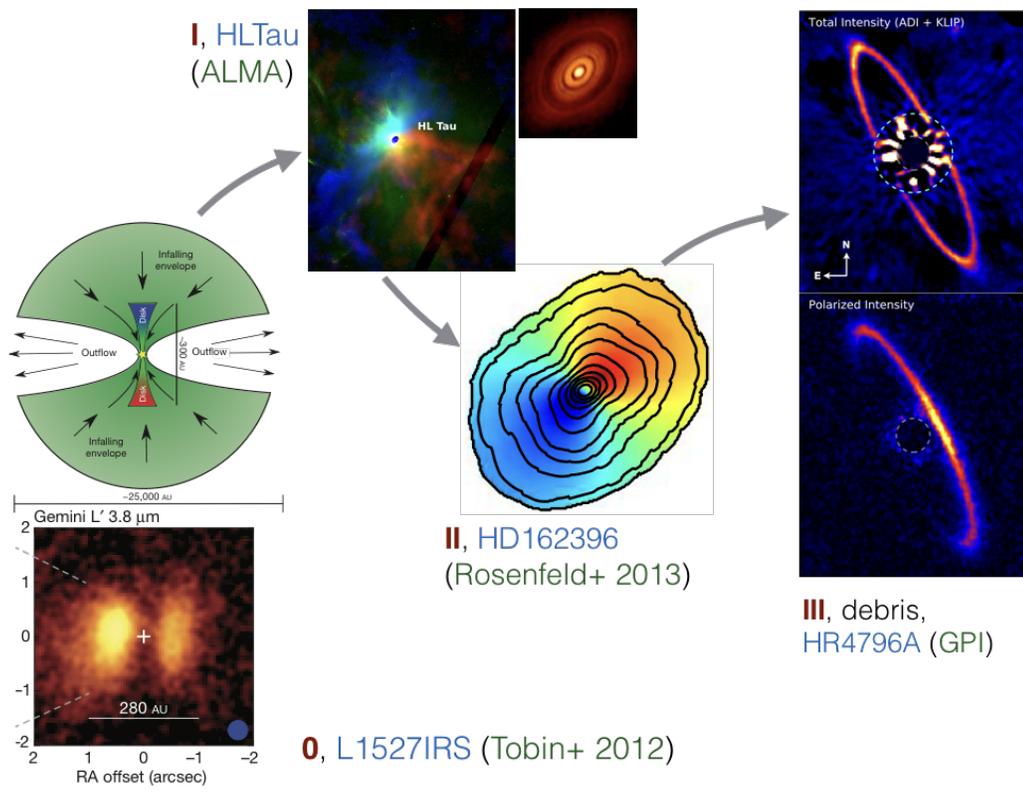


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Example star forming region: Coal sack

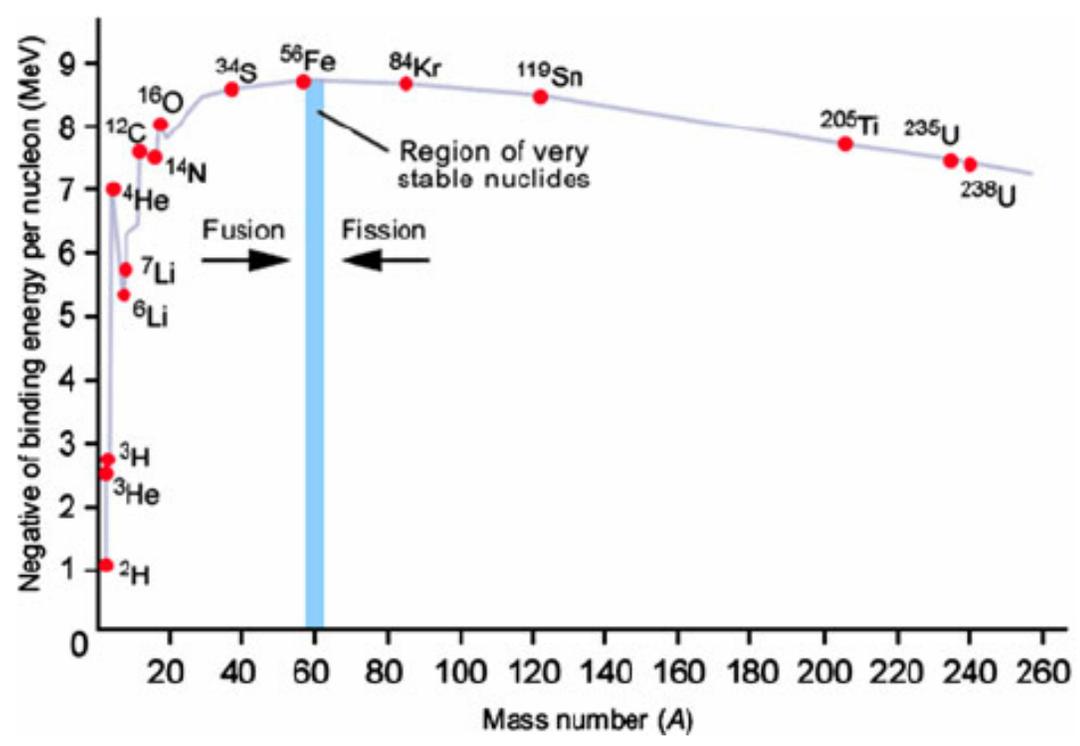


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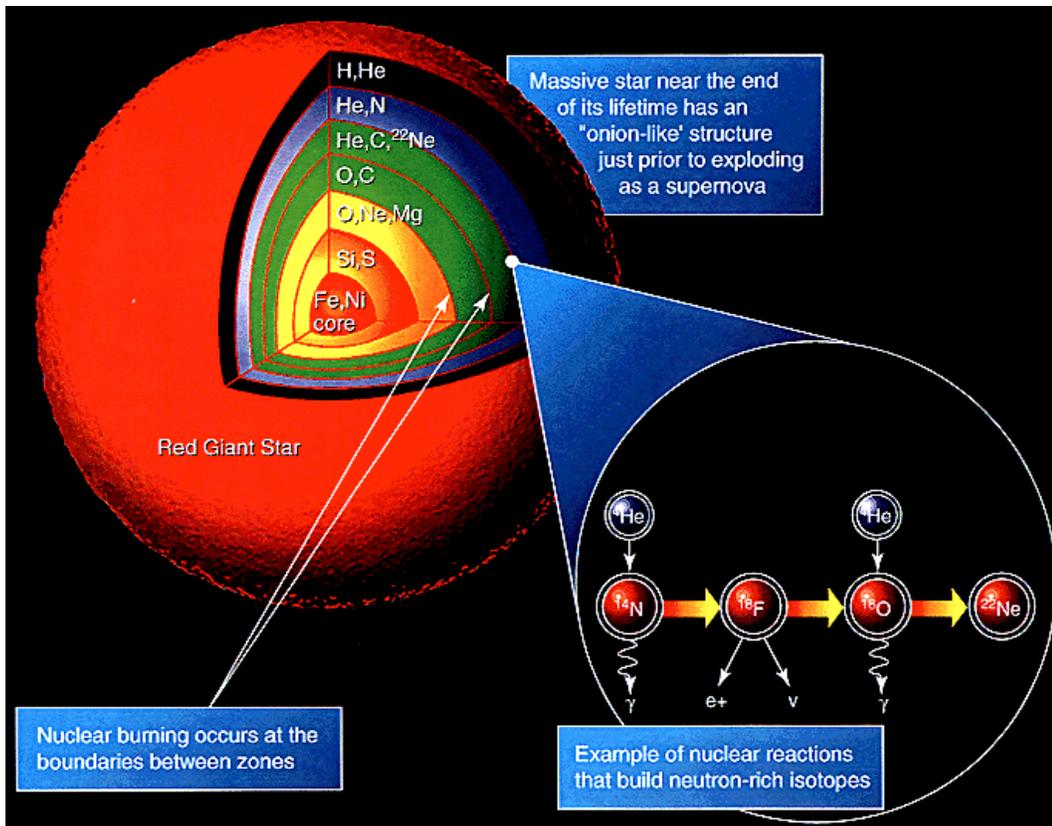


1.3. Nucleosíntesis estelar

Energía nuclear de ligazón



Energía nuclear de ligazón



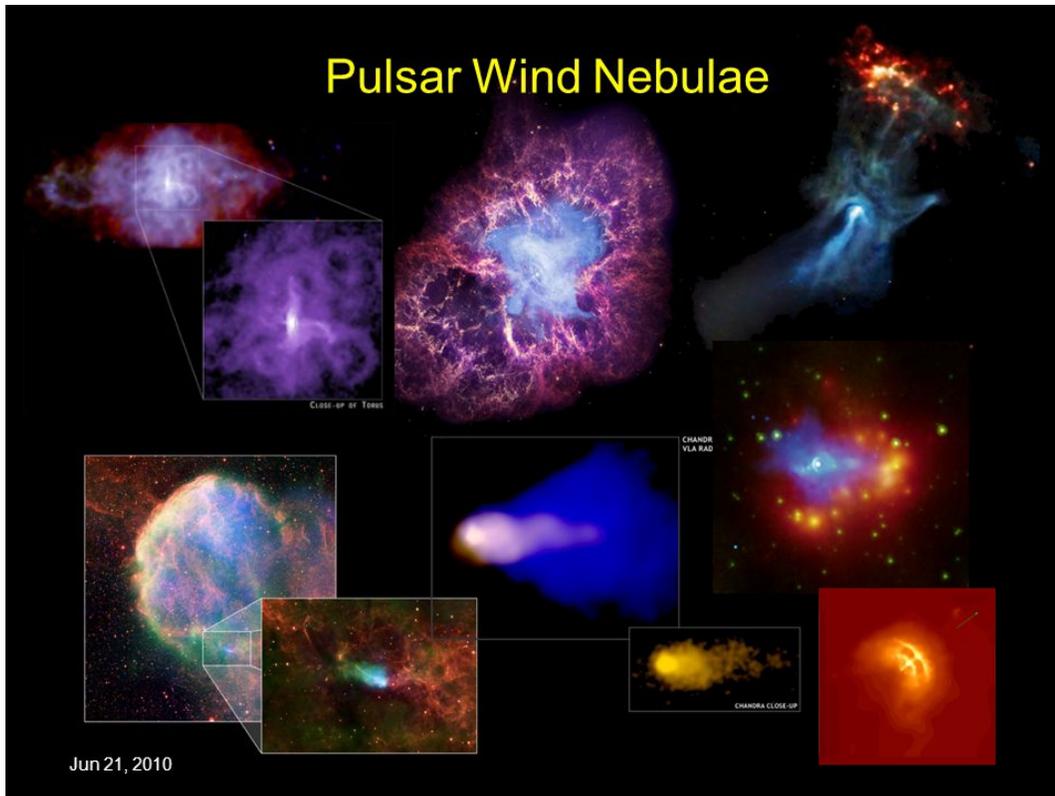
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1.4. Restos estelares

Planetary Nebulae



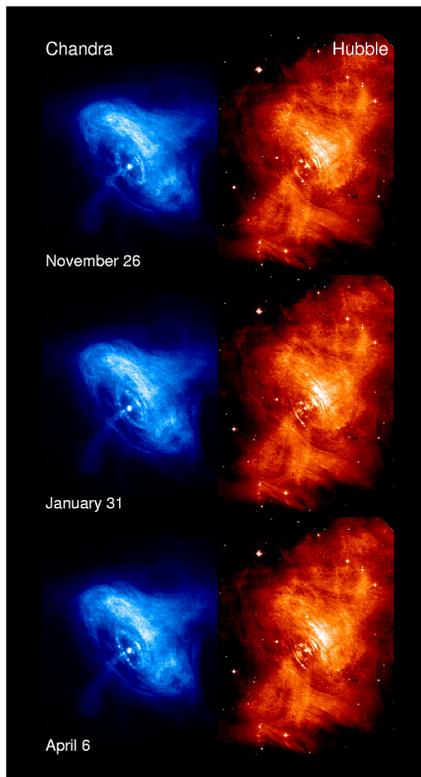
Pulsar Wind Nebulae



Vela Pulsar

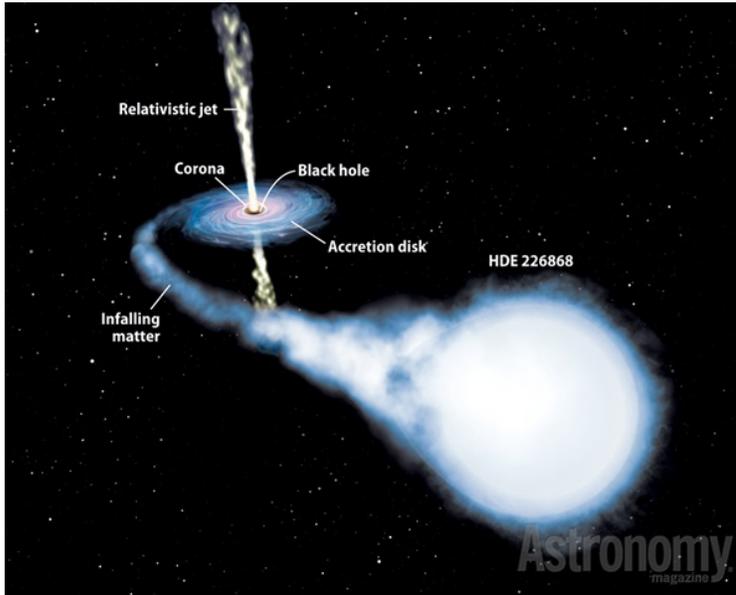


Crab Pulsar



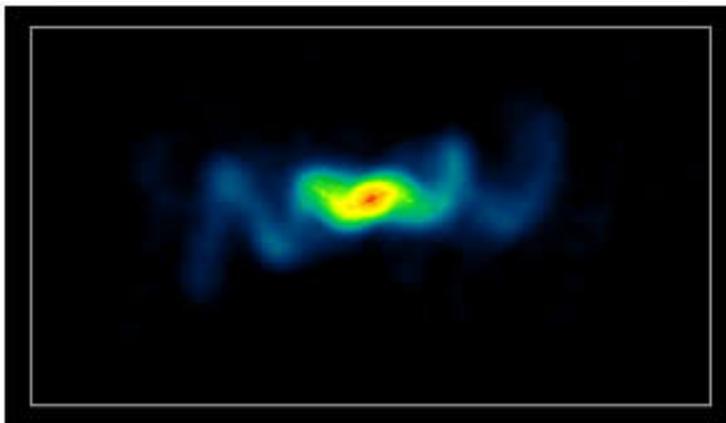
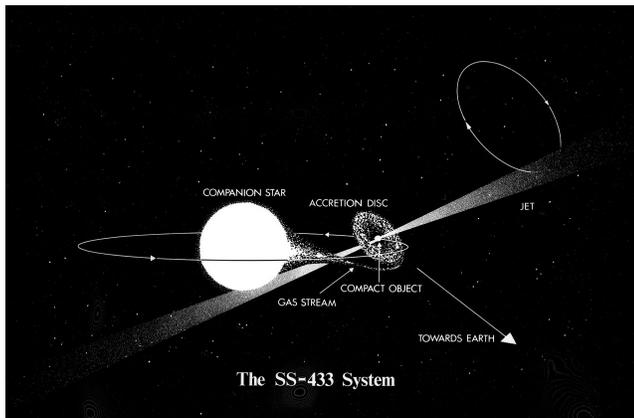
Stellar black holes

- X-ray binaries: estrella + black hole
- más conocidos: SS433, Cyg X-1, GRS 1915+105
- más cercanos GRO J1655-40 o A0620-00 ($\sim 1\text{kpc}$).



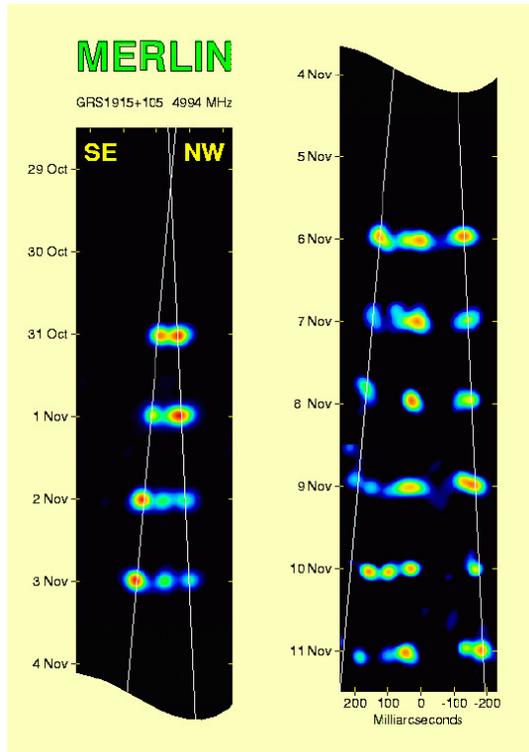
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SS433



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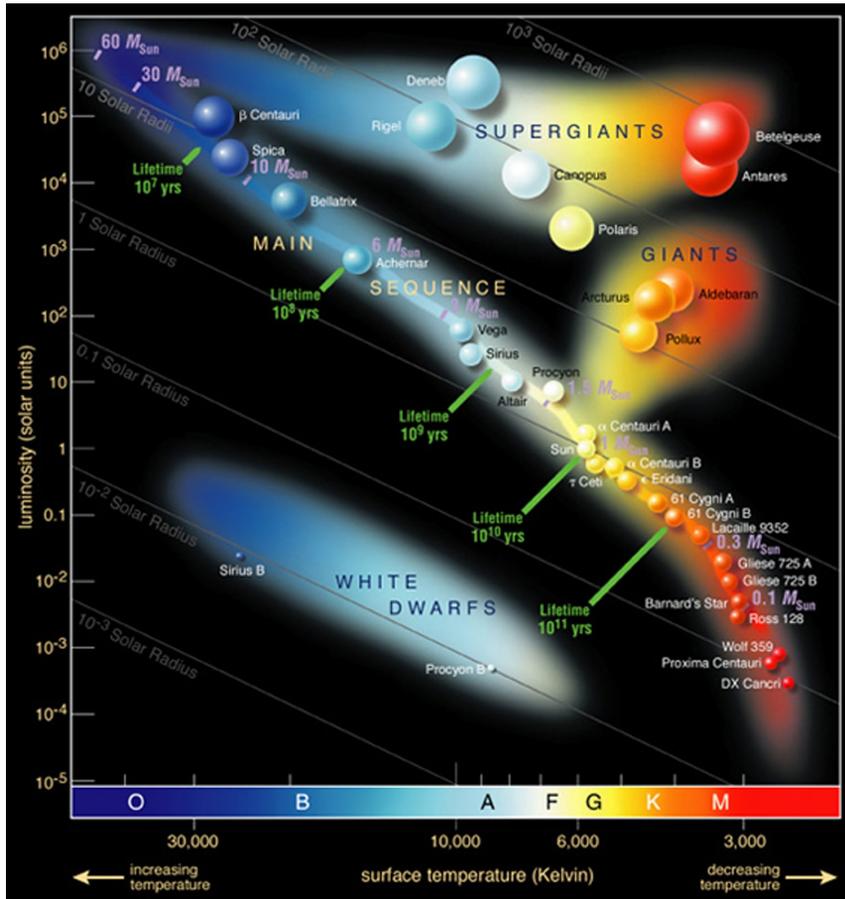
Super luminal jet in GRS1915



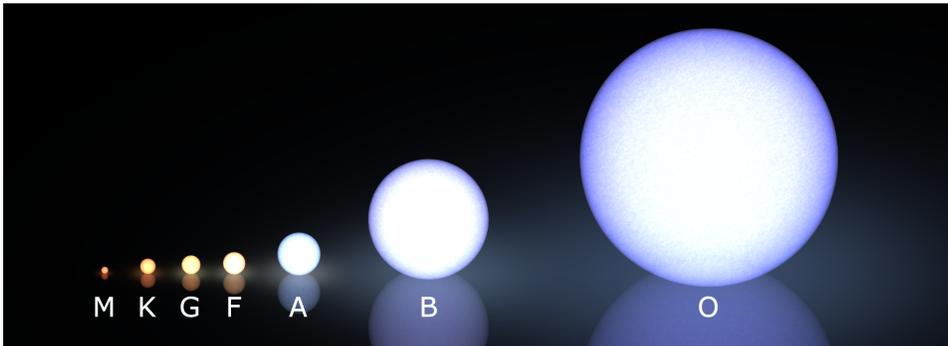
Estrella MS + $\sim 15 M_{\odot}$ black hole.

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1.5. Diagrama H-R



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